LA-UR-16-23112 Status of Cross Section Progress for <sup>235,8</sup>U, <sup>239</sup>Pu, <sup>56</sup>Fe, <sup>16</sup>O

> M.B. Chadwick LANL

(with CIELO collaboration)

WPEC Subgroup 40, OECD, Paris, May 9-13, 2016





#### Abstract

Progress is described for nuclear cross section evaluations, calculations, and experimental measurements at Los Alamos and other laboratories, on 235,8U, 239Pu, 16O and 1H, for the CIELO project at the Nuclear Energy Agency/WPEC. This includes first data from the LANSCE Chi-nu project, providing insights into the energy spectrum of fission neutrons. The net effect of various nuclear data updates is a suite of CIELO files that models criticality well, but feature an improved agreement with differential data (prompt-fission neutron spectra, time of flight experiments in the resonance region, resonance fission neutron multiplicity, quasi-differential neutron scattering data)





## **Overview comments on CIELO progress**

## Progress on our understanding of 1H, 16O, 56Fe, 235U, 238U, and 239Pu neutron reactions:

- a variety of different teams and approaches
- experiments & theory advanced

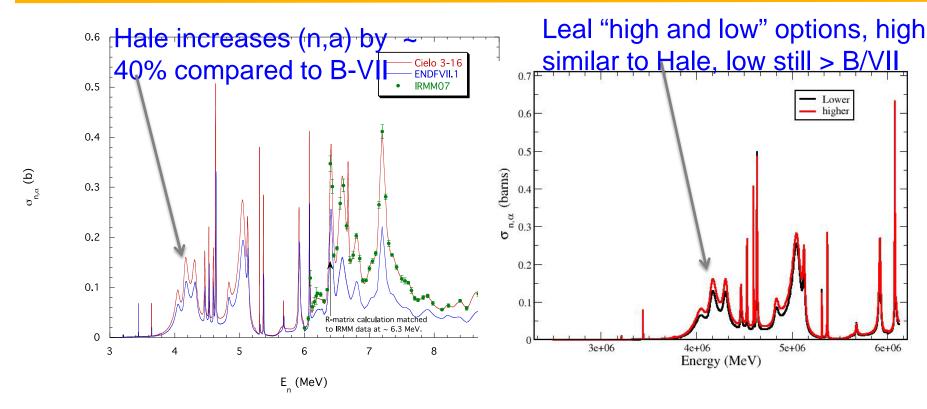
#### A set of starter files has been created via a USA-Europe-IAEA collaboration

- this suite of (ENDF/B-VIII.beta1) files, taken together (with a new H2O scattering kernel), appears to model integral criticality fairly well. Validation testing ongoing (including SG39 support); future refinements planned
- other CIELO collaborations will provide alternative options, e.g. a European JEFF3.3-testing suite (including CEA/BIII+IRSN files) coming.
- Add/refine covariances
- These analyses will be documented in the coming year, including journal articles in Elsevier's Nuclear data Sheets (January 2018)





### Examples of convergence of opinion (1) Oxygen-16 (n,alpha)



Hale file is in accordance with IRMM (Georginis et al.) conclusions

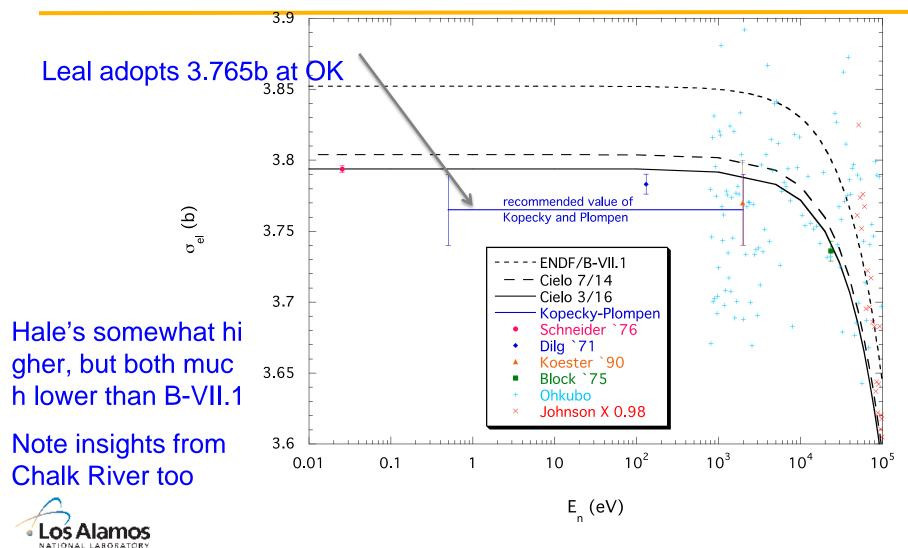
Future "confirmatory" experiments beginning at various labs, including Los Alamos





## **Examples of convergence**

#### **Oxygen-16 low-energy elastic scattering**



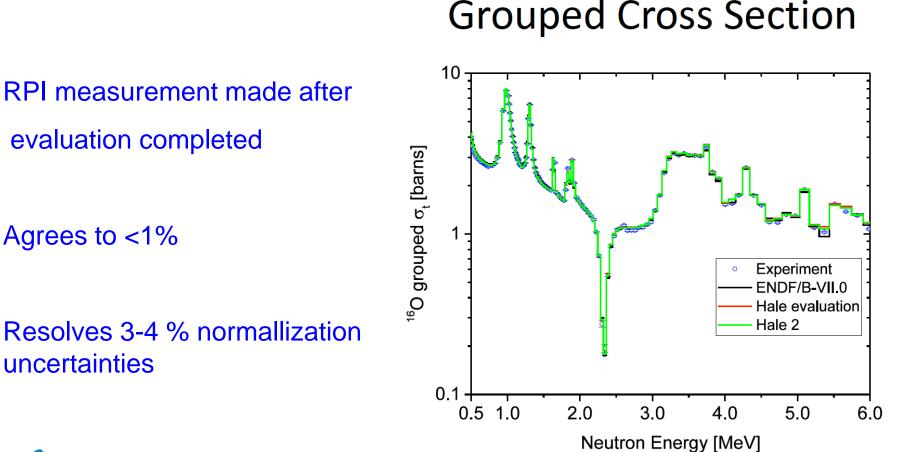
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## **Examples of convergence**

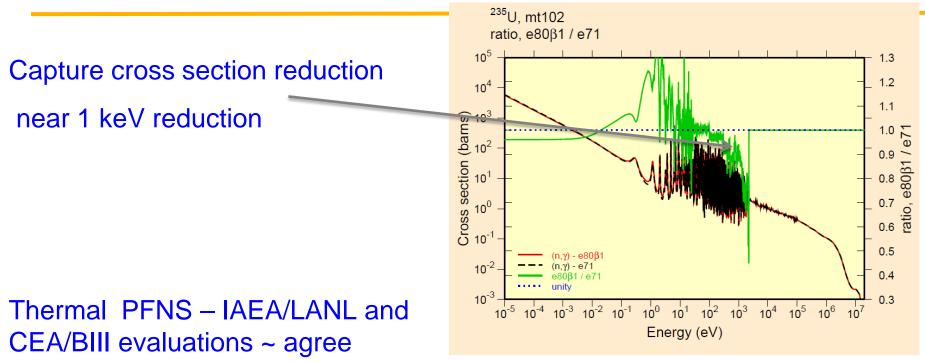
Oxygen-16 total cross section, normalization determined from RPI experiment



Los Alamos
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# Examples of convergence U235



Thermal 235U eav=2.00 MeV

Thermal 239Pu eav~2.10 MeV

(tentative, until we obtain new data)





## **Examples where open questions remain**

- These differences of opinion will be documented

Magnitude of actinide inelastic scattering

Actinide capture

- 238U uncertainties further reduced, largely confirming standard
- 235U significant differences remain in the 10s-100s of keV
- 239Pu new DANCE capture data, but will largely influence
  - a future evaluation (more time needed to include in res analysis)

Various questions in resonance evaluation of 56Fe

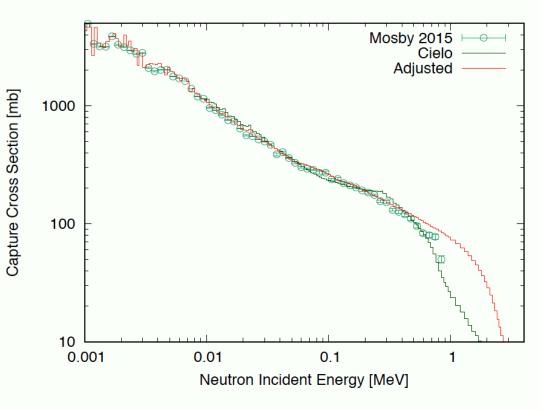




## 239pu n,g

## DANCE Capture Data Test

- Mosby's data Oct. 2015
  - from URR to 1 MeV
  - URR parameters slightly adjusted
  - CoH3 calculation given to the high energy part
    - Soukhovitskii 2005 potential
    - · fission adjusted
    - M1 scissors mode included
- CIELO file issue
  - Inelastic scattering exists in URR, but total does not have it
  - Cross section fixed from 9 to 30 keV



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## **Conclusions on CIELO collaboration**

**CIELO** has stimulated healthy collaborations

Enabled significant progress on these evaluations

- enabled large changes to new regional evaluation files, ENDF, JEFF, ...

We welcome feedback from:

- Integral validation data testers
- Adjustment SG39 project insights
- Sensitivity tools likeNEA's NDaST





## CIELO progress: Scattering (Elastic & Inelastic) for Actinides and Iron

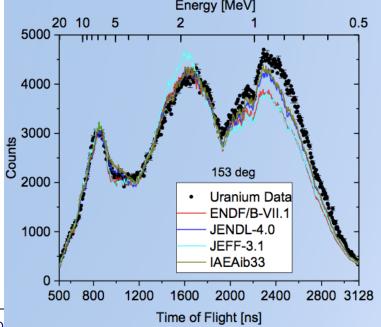
New RPI semi-integral approach has had a large impact on validating and improving inelastic and elastic scattering on 238U and 56Fe

- longer term work on 239Pu and 235U planned at LANSCE

Modern inelastic scattering advances made from theory with insights from measured data:

- 238,5U work from the IAEA
- 235U from BRC
- 239Pu work from LANL
- -56Fe work from BNL, IAEA

## Note SG39 strong sens. to 238U inelastic



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## CIELO progress: Actinide Capture

239Pu: New LANSCE data "first" for many decades (Mosby, Jandel)

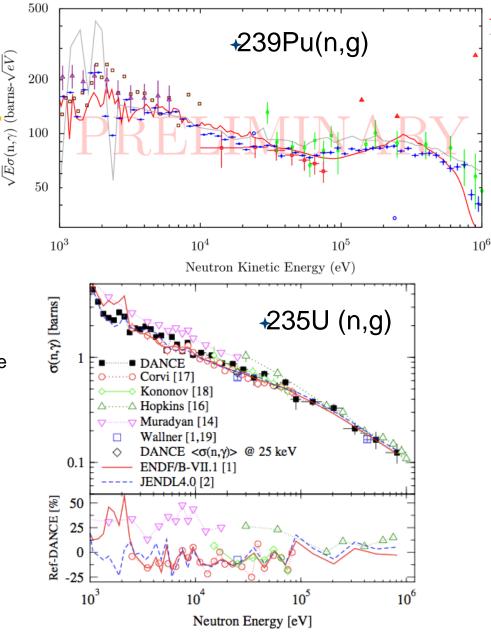
- tend to support ENDF; only modest changes needed, including above 1 keV where we need a SAMMY analyses, as for 235U n,g

235U was a major advance, from 0.5-2 keV (lower capture)

- In tens of keV suggests a small increase to ENDF (but contradicted by Wallner AMS measurement at 25 keV)

238U capture from Schbx

tends to support standards view









IAEA Standards 235U PFNS advances suggest a softer spectrum (2.00 MeV versus 2.03 average energy), but not as soft as Kornilov's 1.97

- concurrent changes to 235U nubar (lowered) being discussed

New LANSCE PFNS from 235U supports "ENDF" Madland-Nix PFNS in the fast range near 2 MeV, as do LANL/NUEX measurements

- This challenges other studies, where softer PFNS are explored

- Leading to a view to not change fast PFNS (much, at least) for 235U and 239Pu, where Chatillon BRC/LANSCE data also ~ supports ENDF. Chinu will measure 239Pu in ~ 1 years. New CIELO file uses Talou-Rising eval. For 235U





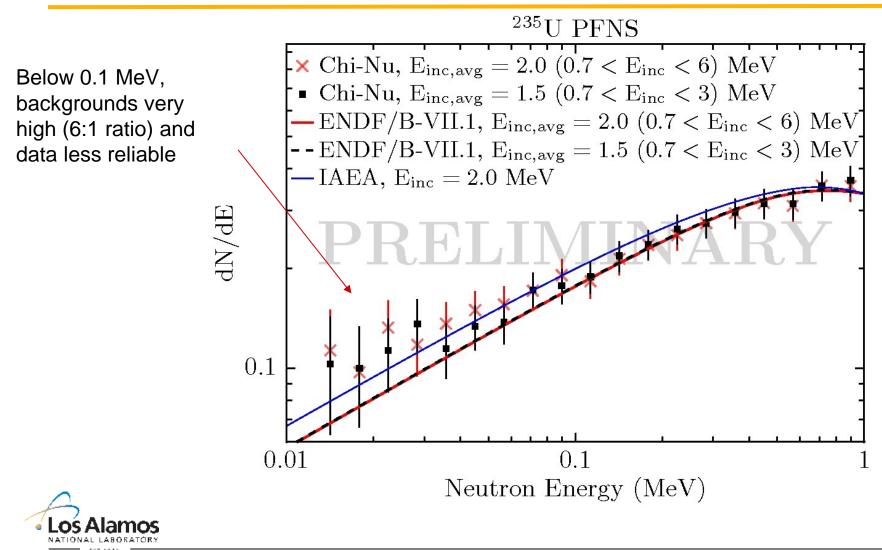
## <sup>235</sup>U: LANL PFNS Experimental Work in Fast Range (0.5-6 MeV) Suggests ENDF PFNS is Accurate

- Lestone data was release last year for 235U as well as 239Pu PFNS (Published in ND2013 proceedings)
  - ℰ Einc average ~ 1.5 MeV
  - ℰ E-emission > 1.5 MeV
- New preliminary data from LANSCE/Chi-nu
  - Einc average various "monoenergetic" and average energy cuts possible, including 1.5 and 2 MeV
  - ℰ E-emission < 1 MeV in first phase of Chi-nu</p>





## <sup>235</sup>U: LANL PFNS Experimental Work in Fast Range (0.5-6 MeV) Suggests ENDF PFNS is Accurate



E31.1945



## <sup>235</sup>U: 2 LANL Experiments cover the whole emission energy range – Chi-nu (LANSCE) and NUEX (Lestone-Shores)

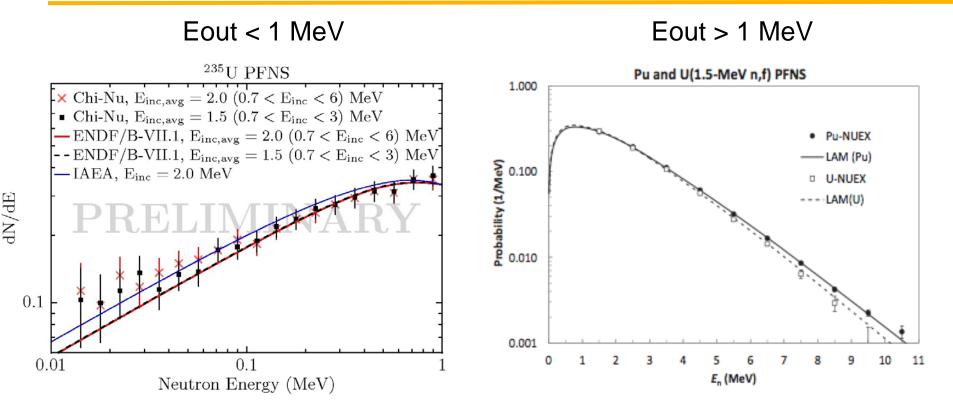


FIG. 3. The emission probabilities listed in Tables III and IV, and the corresponding 1.5-MeV n +  $^{239}$ Pu and  $^{235}$ U Los Alamos fission model fission-neutron energy spectra (curves).





## <sup>239</sup>Pu : Some Particular Challenges

- Build on the excellent WPEC subgroup 34 work from CEA & ORNL
- Capture discrepancies. We're waiting for final DANCE data; preliminary results obtained
- Inelastic scattering discrepancies between evaluations
- Use of new IAEA Standards, including fission (TPC)
- Other new data that will impact the evaluation new PFGS data from DANCE; New FPY data from TUNL (impact esp at 14 MeV)





## Pu-239 Status. Version-0 performs like SG34 at low energies, ENDF/BII.1 at higher energies as expected (See Kahler talk)

#### Contents of the Pu239 file CIELO/B -

- . Based on ENDF/B-VII.1 cross sections
- SG34 resolved resonance parameters
- Prompt nu-bar in JEFF-3.2, up to 650 eV
- Total nu-bar re-calculated
- Base file uses ENDF/B-VII.1 chi< 5 MeV; Romano tweak at thermal: Neudecker>5 MeV [Until we see Chi-nu 239Pu data, we are hesitant to deviate from ENDF in fast range]
- Variants: Other PFNS calculations from Neudecker et al.
- Huge section of delayed gamma-ray spectra removed

#### • Some issues planned to be resolved in this and next years

- Unresolved resonance range, consider use of ISSF = 1 option
- Revise inelastic scattering, in collaboration with CEA/DAM, IAEA, and JAEA
- New gamma-ray production cross section, use of FILE6, and resolve inconsistent fission gamma-ray production
- Upgrade capture cross sections which considers new DANCE data

## Improvements in the new LANL evaluation –

## **Neudecker work**

#### ØExperiment:

Recently published data of Chatillon et al. and Lestone et al. included (+ Granier corretions) Improved uncertainty estimate of exp. data (including Chi-Nu studies)

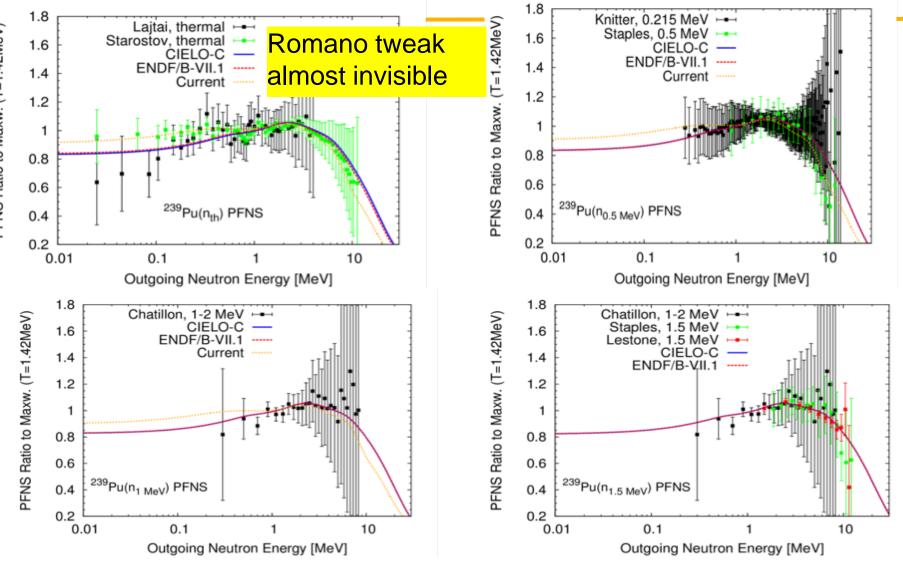
#### ØModeling:

- "Einc-dependent parametrization of <TKE> and <Er> by Lestone et al.
- & Madland was used (constant for ENDF/B-VII.1)
- "Pre-equilibrium component of the PFNS considered via CoH
- "Only neutrons coming from the fission process are counted

#### ØEvaluated output:

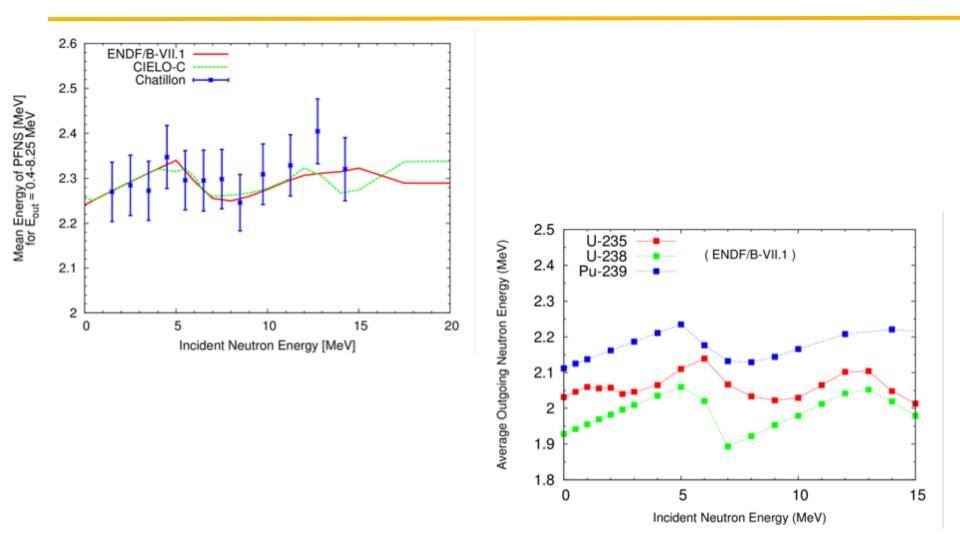
- $_{u}$ Given for Einc= thermal 30 MeV
- "Evaluated covariances are given for all Einc and also between different Einc

# 239Pu PFNS at all given Einc, Compared to Neudecker's Current Evaluation

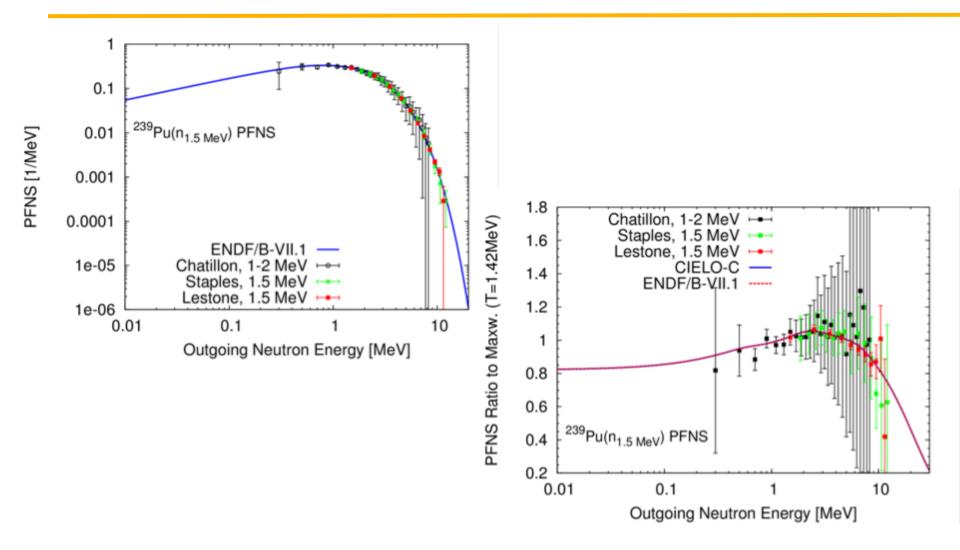


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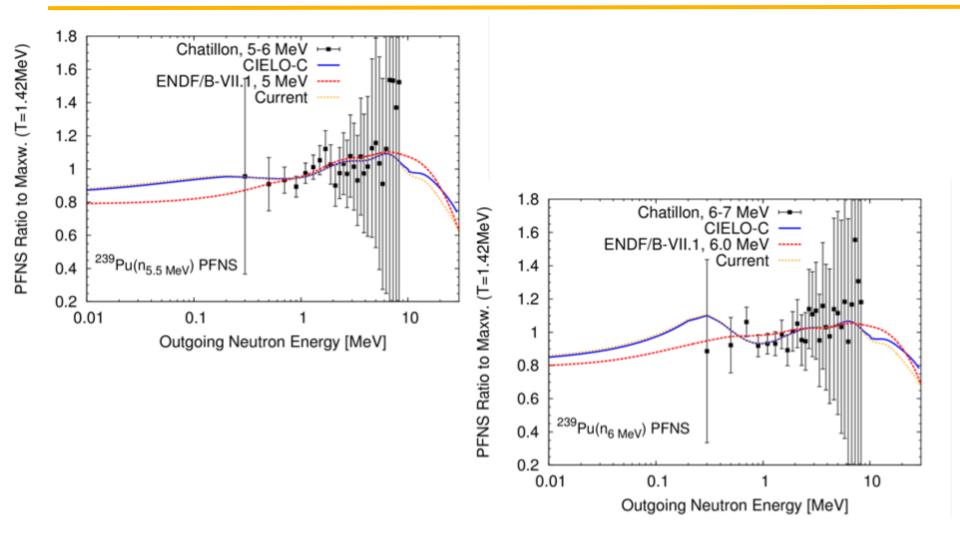
## PFNS Average Energy – CIELO file for testing (ENDF <5 MeV except for a tweak at thermal by Romano, and Neudecker > 5 MeV)



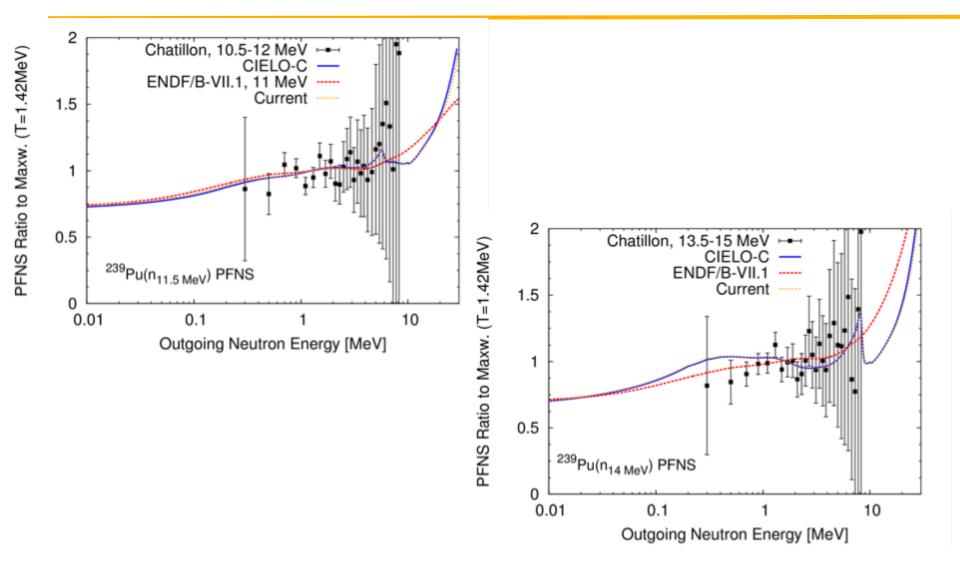
## +239Pu PFNS at Einc = 1.5 MeV



# +239Pu PFNS at Einc = 5.5-6 MeV (opening of second chance fission)



## +239Pu PFNS at Einc = 11.5 MeV and Einc = 14 MeV

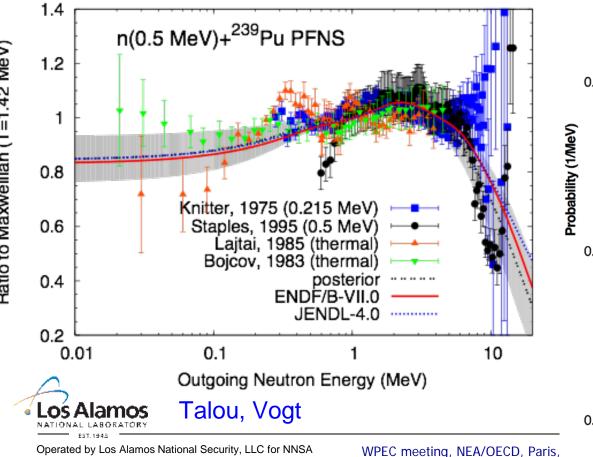


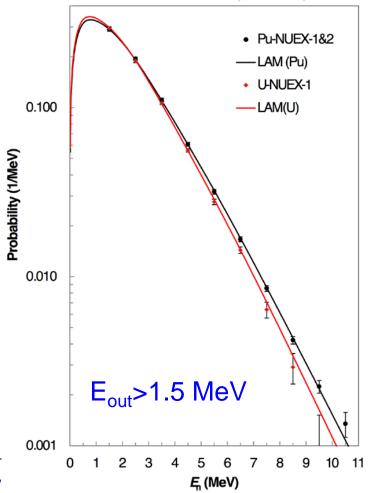
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## **Determining the Prompt Fission Neutron Spectrum (Chi): One of Our Highest Priorities & an IAEA CRP.** Chi-nu PFNS delayed till next year (235U measured recently)

Large uncertainties below 1 MeV and above 5 MeV impact criticality calculations and (n,2n) transmutations Lestone's talk: accurate underground NUEX data released by Los Alamos:

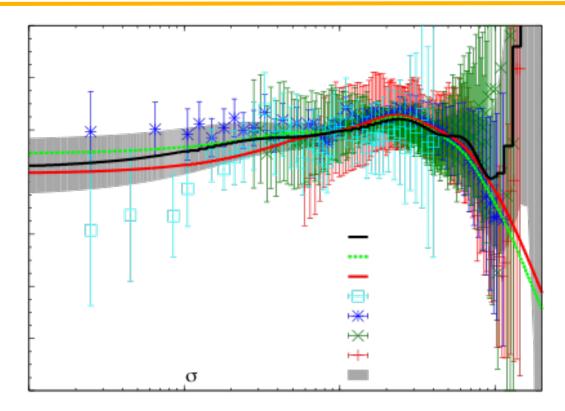
Pu-NUEX-1&2 and U-NUEX-1 (1.5-MeV n,f) PFNS





Ratio to Maxwellian (T=1.42 MeV)

## Ongoing work on PFNS by Talou, Reisner, Neudecker (red = cielo.0 ; green = cielo.1)



Black = snapshot of ongoing work. Will be updated to include Lestone, Chatillon, *etc* 

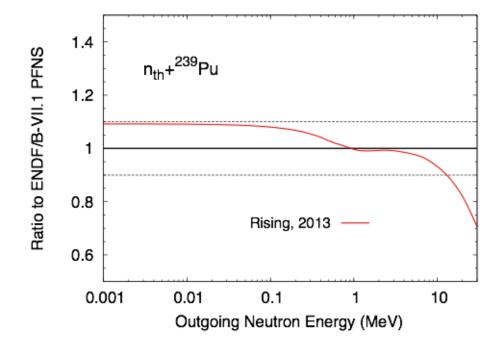
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# Ongoing work on PFNS – Reisner result for thermal, in file cielo.1 for testing







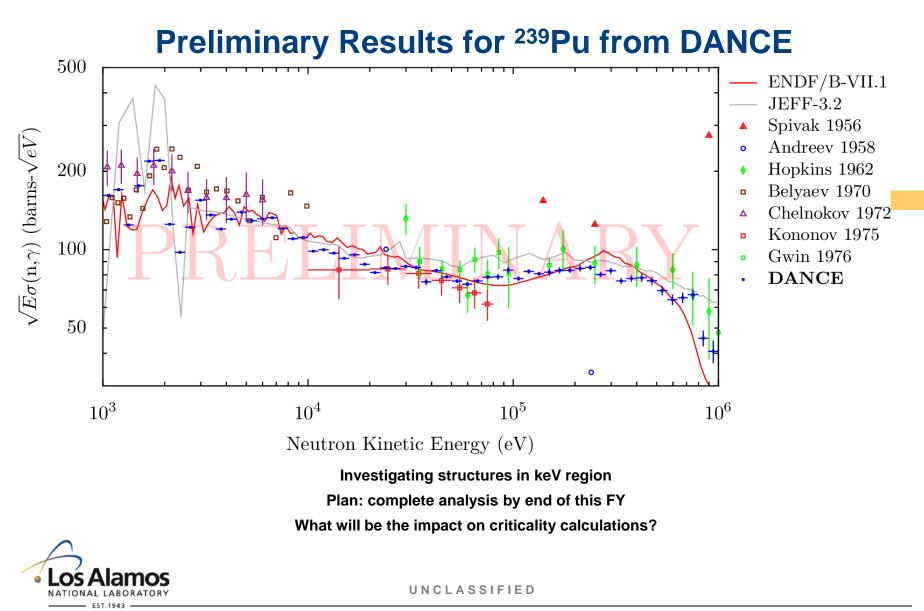
### **Plutonium Capture: Improvements Are Needed**

## **Existing uncertainties >15%**

+SG33 & PROFIL (PHENIX) <sup>239</sup>Pu(n,γ) 239Pu(n,g) integral testing suggests 10<sup>1</sup> Cross Section B-VII is ~ 10% low over this fast Section (b) reactor spectrum. Also, Ishikawa's 10<sup>0</sup> ADJ work suggests JENDL should be raised 5-10% DANCE measurements now being 10 Cross ENDF/B-VII.1 analyzed ENDF/B-VII JEFF-3.1 JENDL-3.3 10<sup>-2</sup> Kononov, 1975 Gwin, 1976 Hopkins, 1962 & Ben Diven **DANCE**, 2011 10<sup>-3</sup>` 10<sup>-2</sup> 10<sup>-1</sup> -3 10<sup>0</sup> 10 Neutron Energy (MeV)

EST.19

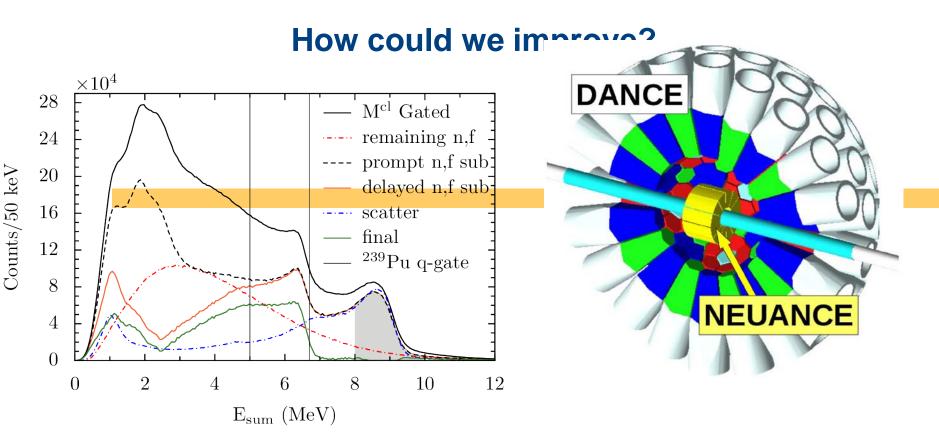




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Fission and scattered neutron background strong above 10 keV (left) Neutron detector inside DANCE (right) could reject much of this Prototype detector run in January – optimizations are needed



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#### **Inelastic Scattering Discrepancy**

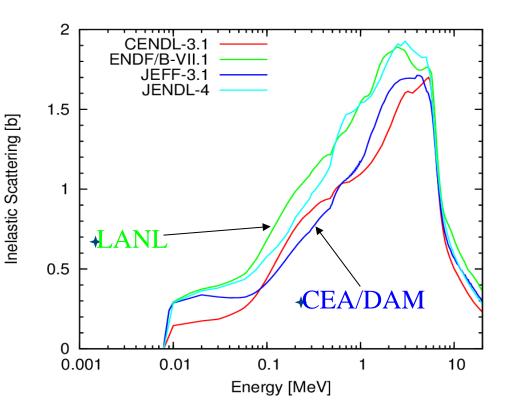
#### IAEA Technical Meeting on Model Calculation for Major Actinides

• Summary report published: INDC(NDS)-0597, R. Capote, et al.

+These two files equally work for Jezebel keff prediction.

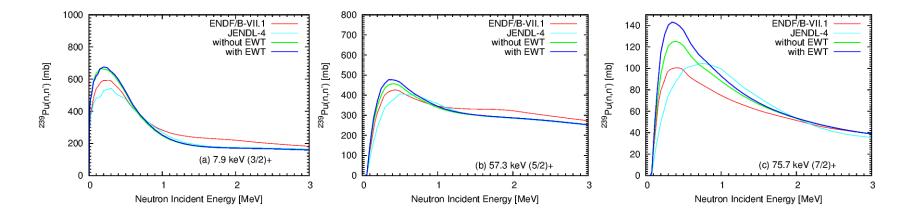
Probably, the difference in the inelastic scattering comes from the optical potential parameters adopted in each library

- CEA total cross section is higher than ENDF in the 30keV - 500keV range
- total and absorption cross sections anti-correlated

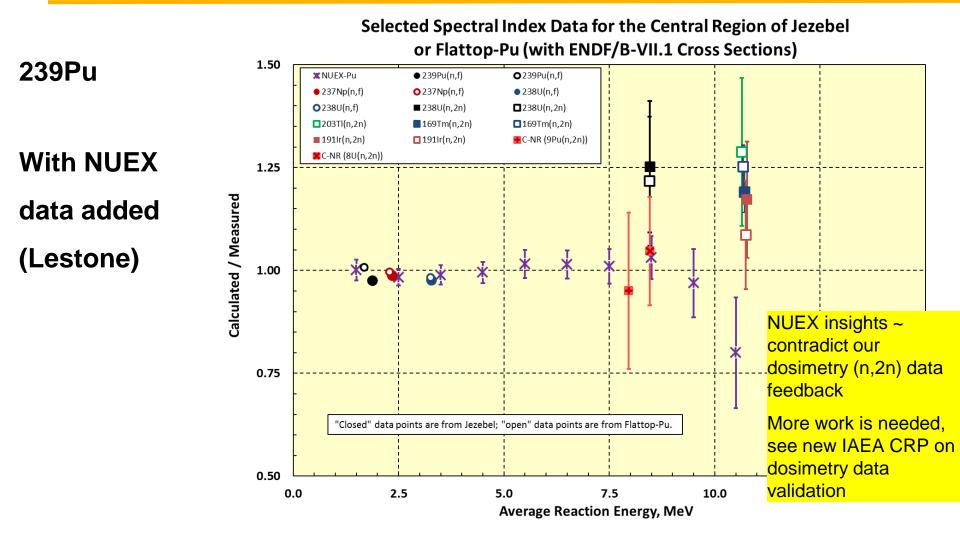


### **Pu-239 Inelastic Scattering - Kawano and collaborators**

- Correct treatment of compound cross section
  - . Full Engelbrecht-Weidenmueller (EW) transformation performed
  - Fission channel has not yet optimized
    - higher than evaluations
  - . Difference between the EW and Hauser-Feshbach-Moldauer cases seems to be small



## Reaction Rates in Fast Critical Assemblies Provide Integral Test of Prompt Fission Neutron Spectrum & (n,2n) Cross Sections - *Plutonium-239 PFNS Data*



WPEC meeting, NEA/OECD, Paris, May 9, 2016

## New <sup>16</sup>O Evaluation Based on R-Matrix Analysis of the <sup>17</sup>O System

### G. M. Hale and M. W. Paris, T2



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WPEC meeting, NEA/OECD, Paris, May 18-22, 2018 - 16-22419

## **Major advances for 160**

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- Higher (n,a) cross sections informed by both Georginis re-analysis of older data, and by constraints from scattering theory
- Lower total/elastic low-energy cross section
  - not quite as low as Kopecky-Plompen analysis, but lower than B-VII
  - Total cross section validated though recent RPI measurement
  - Confirmation data, obtained after evaluation was completed, agrees to 1%

## Outline

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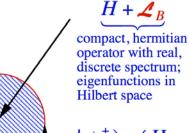
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- Reminder of R-matrix properties, EDA code
- Status of the <sup>17</sup>O system analysis and <sup>16</sup>O evaluation
  - Low-energy scattering cross sections
  - **1**<sup>3</sup>C( $\alpha$ ,n) and <sup>16</sup>O(n, $\alpha_0$ ) cross sections
  - Fits, data renormalizations, etc.
  - Extension of the evaluation to higher energies
- Summary and conclusions

#### **R-matrix Formalism**

INTERIOR (Many-Body) REGION (Microscopic Calculations)



(S-matrix, phase shifts, etc.)  $(r_{c'} | \psi_c^+ \rangle = -I_{c'}(r_{c'})\delta_{c'c} + O_{c'}(r_{c'})S_{c'c}$ 

or equivalently,

$$(r_{c'} | \psi_c^+ \rangle = F_{c'}(r_{c'}) \delta_{c'c} + O_{c'}(r_{c'}) T_{c'c}$$

**ASYMPTOTIC REGION** 

 $|\psi^+\rangle = (H + \mathcal{L}_B - E)^{-1} \mathcal{L}_B |\psi^+\rangle$ 

Measurements

SURFACE  

$$\mathcal{L}_{B} = \sum_{c} |c| (d \left( \frac{\partial}{\partial r_{c}} r_{c} - B_{c} \right),$$

$$(\mathbf{r}_{c} | c) = \frac{\hbar}{\sqrt{2\mu_{c}a_{c}}} \frac{\delta(r_{c} - a_{c})}{r_{c}} [(\phi_{s_{1}}^{\mu_{1}} \otimes \phi_{s_{2}}^{\mu_{2}})_{s}^{\mu} \otimes Y_{l}^{m}(\hat{\mathbf{r}}_{c})]_{J}^{M}$$

$$R_{c'c} = (c' | (H + \mathcal{L}_{B} - E)^{-1} | c) = \sum_{\lambda} \frac{(c' | \lambda)(\lambda | c)}{E_{\lambda} - E}$$

#### R-Matrix Analysis of Reactions in the <sup>17</sup>O System

	channel	nnel a <sub>c</sub>		I <sub>max</sub>		
	n+ <sup>16</sup> O	4	4.4	4		
	α+ <sup>13</sup> C	ļ	5.4	5		
Reaction	Energies (MeV)		# dat point			Data types
<sup>16</sup> O(n,n) <sup>16</sup> O	$E_{n} = 0 - 7$		254	0	σ	$T_{T}, \sigma(\theta), P_{n}(\theta)$
<sup>16</sup> O(n,α) <sup>13</sup> C	$E_n = 2.35 -$	5	67	2	$\sigma_{i}$	nt, σ(θ), $A_n(\theta)$
<sup>13</sup> C(α,n) <sup>16</sup> O	$E_{\alpha} = 0 - 5.4$	1	87	0		$\sigma_{int}$
$^{13}C(\alpha,\alpha)^{13}C$	$E_{\alpha} = 2 - 5.7$	7	1168			σ(θ)
total			525	0		8

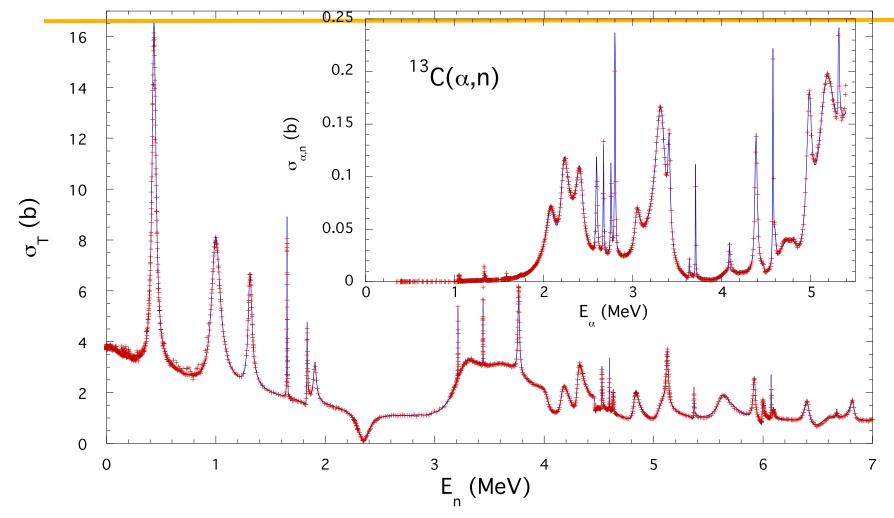
 $-\chi^2$  per degree of freedom = 1.68

#### **Total Cross Section Data**

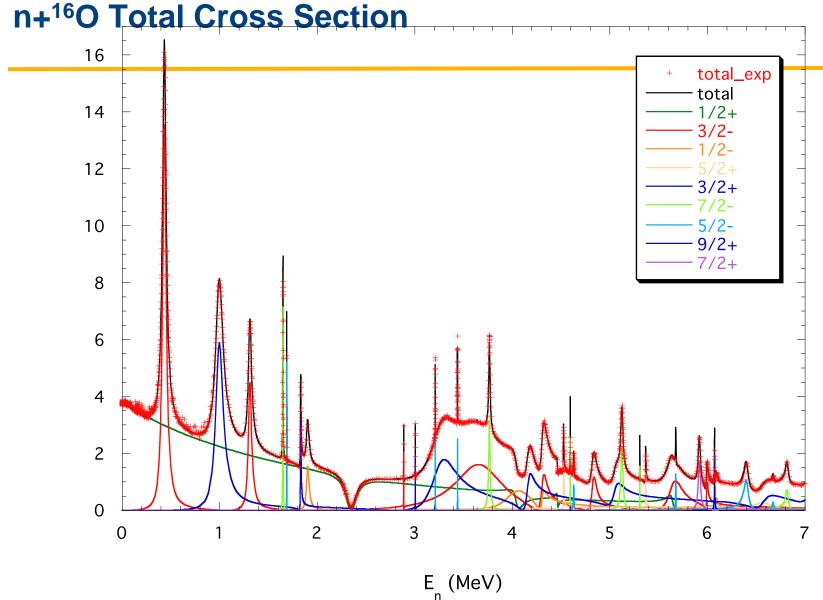
Authors (n,n):	Energy Range	-Energy Shift	Normalizati on
+Schneider	<b>→</b> 0.0253 eV	+0	+1.0 (fixed)
Dilg,Koester,Bl ock	<b>√</b> 0.13 – 23.5 keV	+0	+1.0 (fixed)
Ohkubo (corr. for H)	40.8 – 935 keV	+0	<b>4</b> 0.9989
Johnson &	<b>4</b> 9 − 3139 keV	+0	+0.9799
-Authors (α,n):	Energy Range	Energy Shift	Normalizati on
Drotleff et al.	<b>√</b> 346 – 1389 keV	+0	+1.0 (fixed)
<b>.</b> Heil et al.	<b>√</b> 416–899 keV	+0	1.0 (fixed)
<b></b> ⊀Kellogg	+445–1045 keV	+0	+1.506
Bair and Haas	<b>√</b> 0.997–5.402 MeV	≁4 keV	<b>-</b> 0.9410

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#### **Integrated (total) Cross Sections**



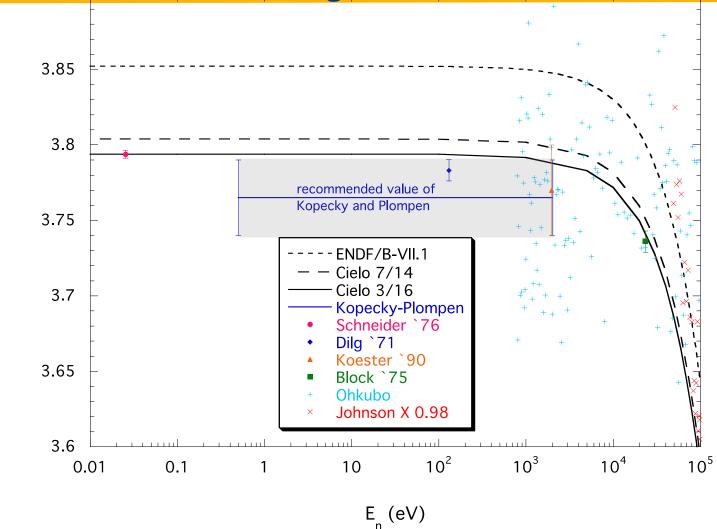
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 $\sigma_{T}^{}(b)$ 

#### n+<sup>16</sup>Q<sub>9</sub>Elastic Scattering Cross Section



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 $\sigma_{el}^{}\left( b\right)$ 

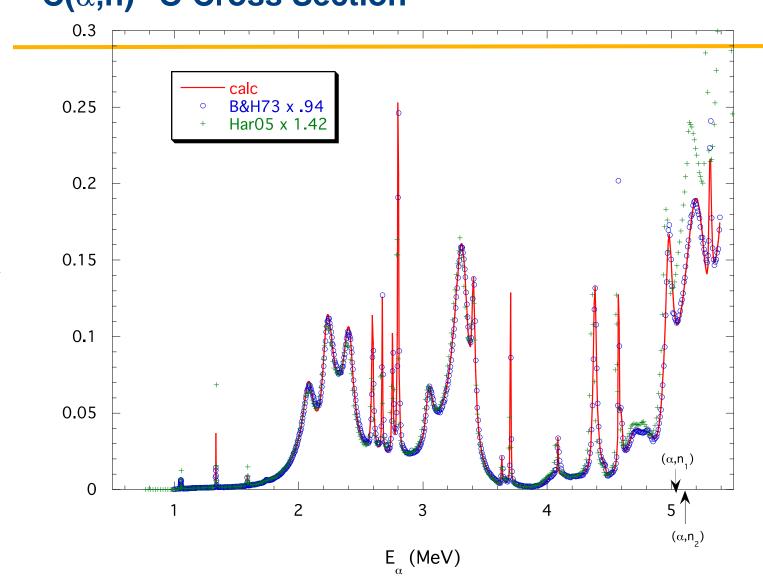
#### Giorginis' Analysis of ( $\alpha$ ,n) Measurements

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- Considered two measurements, Bair and Haas (B&H73) and Harissopulos *et al.* (Har05).
  - Determined a preliminary cross-section scale for B&H73 based on the integral of the thick-target yield over the narrow resonance at 1.056 MeV that agrees with the published scale of Har05.
    - Then applied a correction common to both data sets related to characterization of the <sup>13</sup>C target that gives the cross-section scales 0.95×B&H73 and ~1.42×Har05.
    - Considers the relative shape of the B&H73 measurement to be the most accurate since it had the thinnest target.



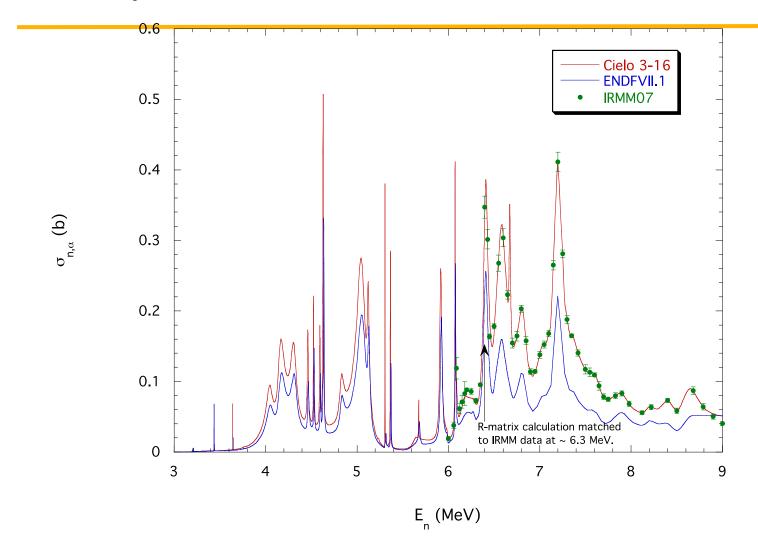
<sup>13</sup>C(α,n)<sup>16</sup>O Cross Section

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+UNCLA +4

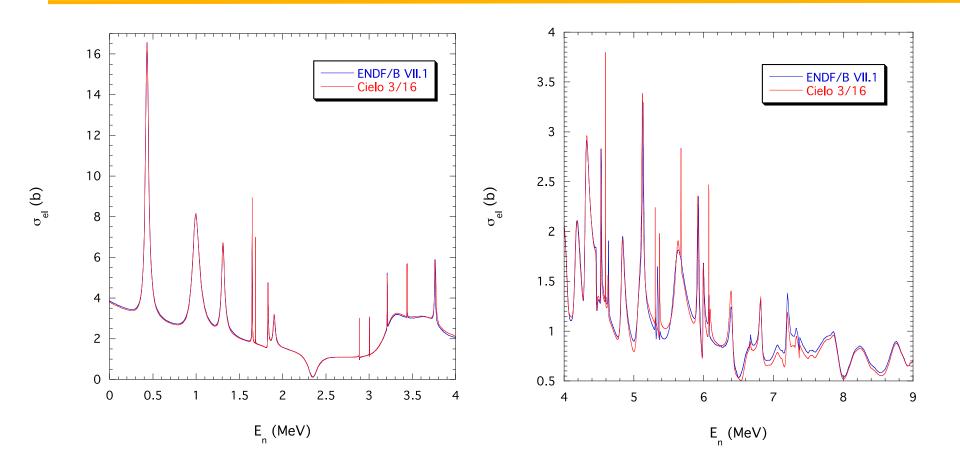
 $\sigma_{\alpha,n}(b)$ 

### <sup>16</sup>O(n, $\alpha_0$ )<sup>13</sup>C Cross Section



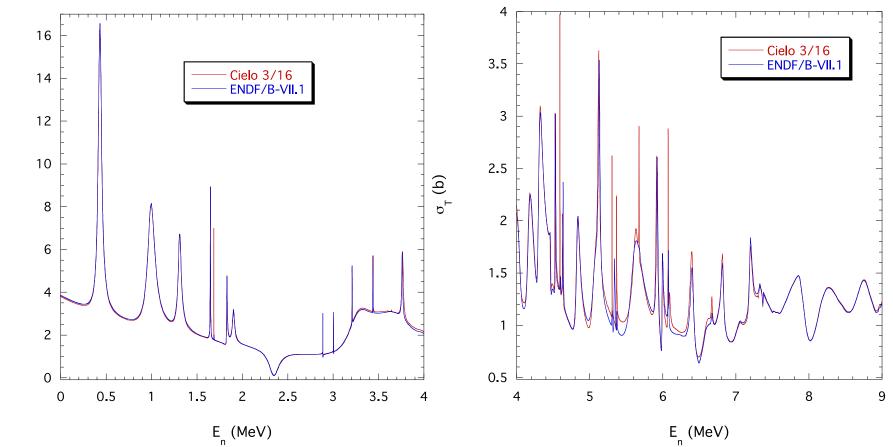
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#### n+<sup>16</sup>O Elastic Cross Section



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#### n+<sup>16</sup>O Total Cross Section

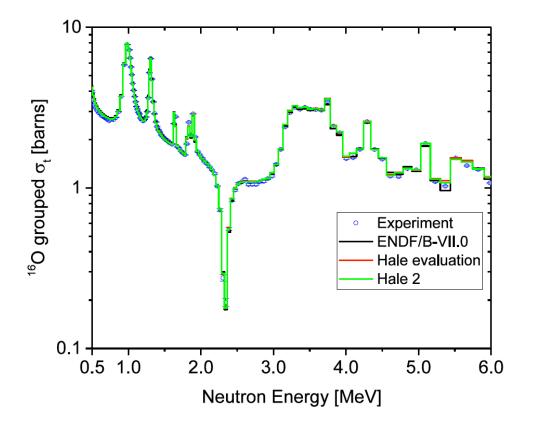


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 $\sigma_{_{T}}\left( b\right)$ 

n+<sup>16</sup>O Total Cross Section – Comparison to RPI: Integral in 3.2-6 MeV region, C/E=1.005 +/- 0.003. (Confirmatory data, not included in Hale's analysis),





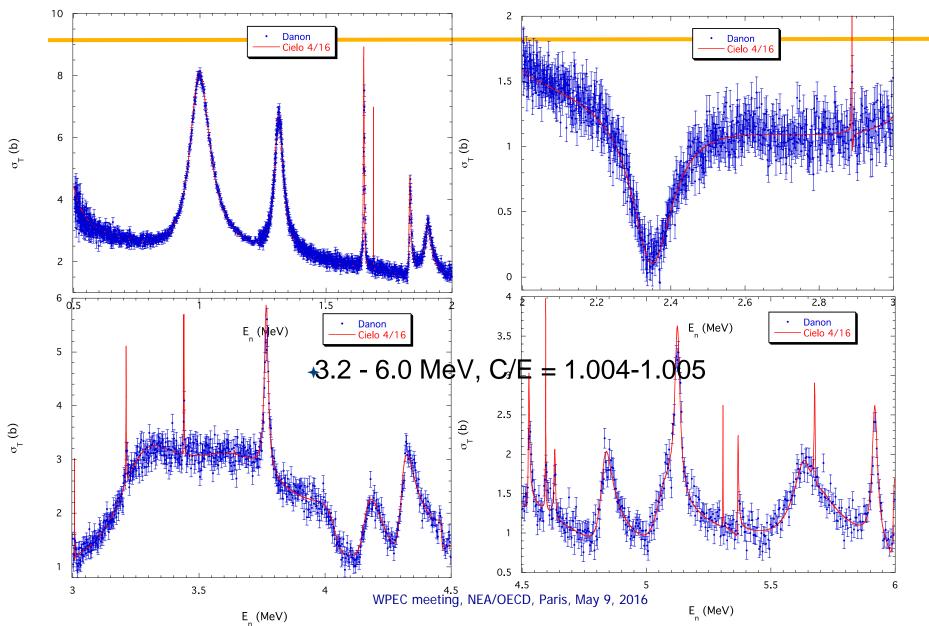
Hale estimate of his

evaluated uncertainties:

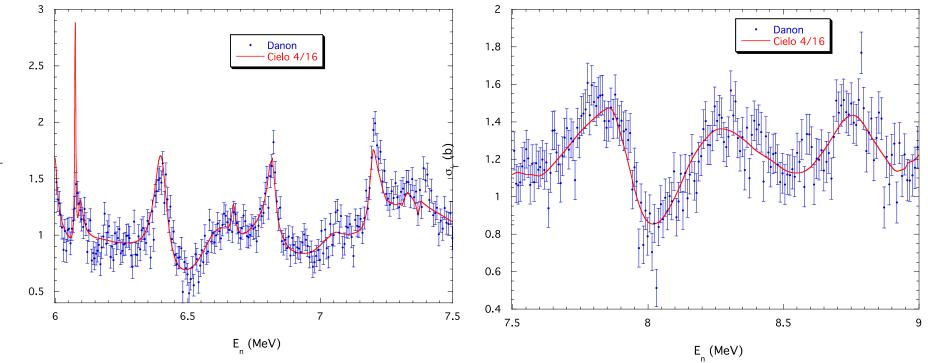
- 0.5 2.0 MeV: 1.99 %
- 2.0 3.2 MeV: 3.03 %
- 3.2 6.0 MeV: 2.60 %
- 6.0 9.0 MeV: 7.59 %

# evaluation

0.5 - 2.0 MeV, C/E = 0.996 2.0 - 3.2 MeV, C/E = 0.988-0.990

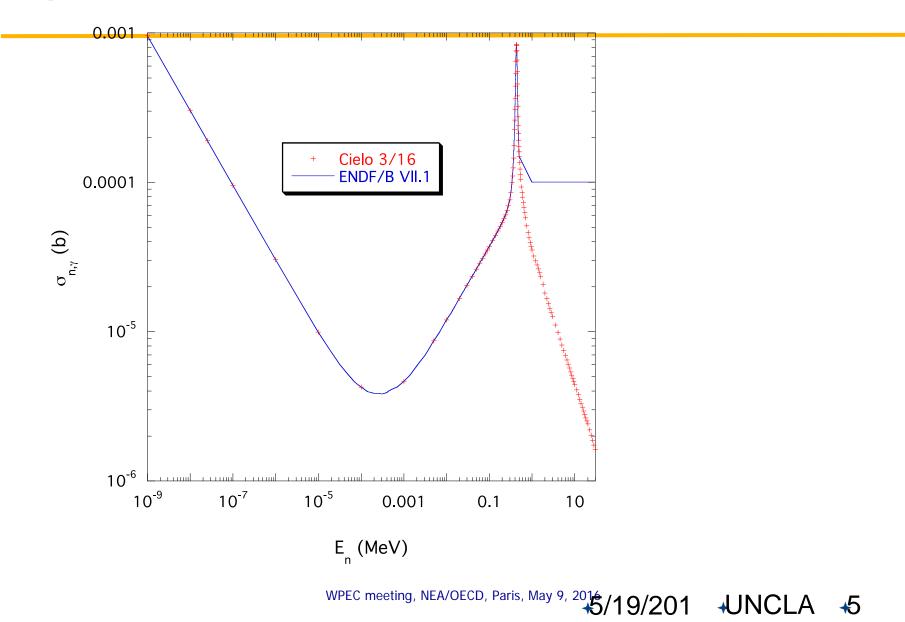


**Comparison, cont.** +6.0 - 9.0 MeV, C/E = 0.992



 $\sigma_{_{T}}^{}(b)$ 

#### **Capture Cross Section**



#### **Summary and Conclusions**

- R-matrix descriptions are constrained by fundamental properties (unitarity, causality, TRI) of nuclear reaction theory.
- EDA analysis of the <sup>17</sup>O system includes data from all possible reactions, giving results that are highly constrained by the properties above (especially unitarity).
- The low-energy n+<sup>16</sup>O scattering cross sections are now in better agreement with high-precision measurements, and the (n, $\alpha_0$ ) cross section agrees with the data of B&H73, IRMM07 (Giorginis).
- The evaluated <sup>16</sup>O file Cielo 3/16 extends to 150 MeV, and is the same as ENDF/B VII.1 above 9 MeV (except for capture).

### 56Fe iron evaluation advances - BNL, IAEA, ORNL, IRNS

- New evaluation makes use of most precise and most recent experimental data between resonance region and 4 MeV (Berthold data for total, new Geel (Negret, Plompen) data for inelastic, and old but not used before data by Dupont for inelastic, Kinney data for elastic angular distributions)

- It uses IRDFF file for the dosimetry reaction (n,p) (calculations agree with IRDFF within uncertainties except of low incident energies)
- Employs dispersive, Lane-consistent optical model above 4 MeV
- Uses consistent modeling up to 150 MeV
- Is informed by the semi-integral data from RPI (neutron emission and capture)
- Provides better reproduction of inelastic (low energies by construction, higher energies by improved modeling)





#### 56Fe iron evaluation advances - BNL, IAEA, ORNL, IRNS

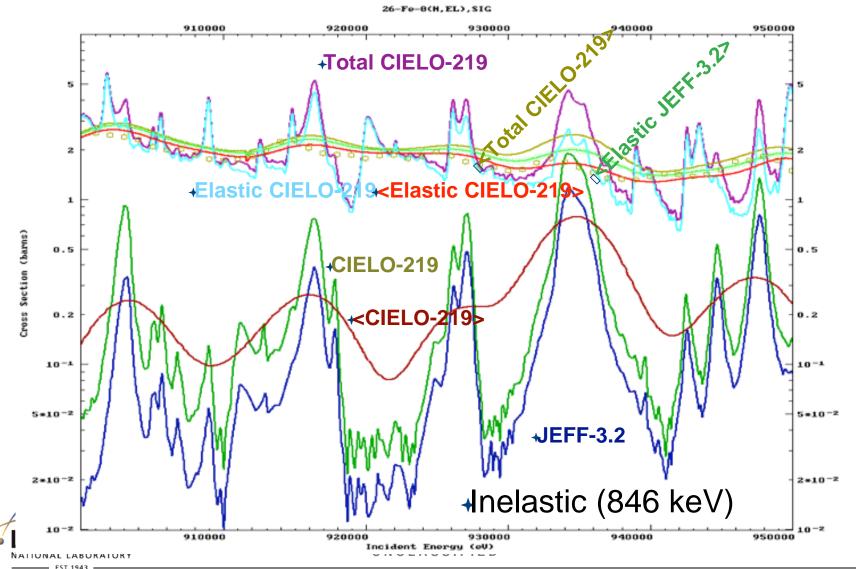
- Reproduces experimental neutron spectra better than in VII.1

- Improves agreement for the criticality benchmarks by adjusting angular distributions in the resonance region and adding background to capture to simulate the effect of lost d-wave resonances (we'll work on resonance region to remove backgrounds and angular distribution tweaks while maintain benchmark results)





#### nat-Fe: Total, Elastic, Inelastic 900-950 keV



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WPEC meeting, NEA/OECD, Paris, May 18-22, 2015



# **Elastic angular distributions**

Kinney data are the most extensive and detailed above the inelastic threshold

JEF-2.2=>JEFF-3.2 ang. distr. are fitted to the Kinney data

Whenever low energy-resolution experimental data are available they are closer to EMPIRE than to Kinney

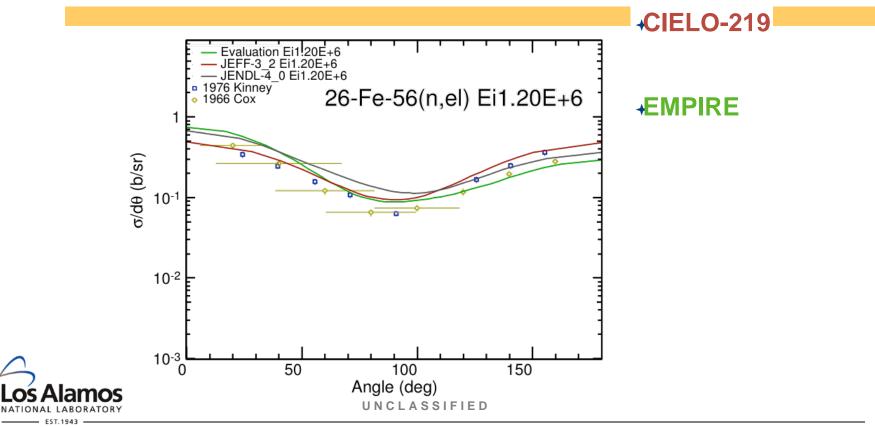
However, RPI semi-integral experiment favors JEF(F)s so we adopted it between 846 keV and 4 MeV

RPI broad-average data compared with EMPIRE and broad-averaged

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#### **Elastic angular distributions – Kinney data**



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WPEC meeting, NEA/OECD, Paris, May 18-22, 2015



### 235U iron evaluation advances - IAEA, ORNL, IRNS, LANL & Performance in Criticality Simulations – Low energy RR

fit of Thermal Neutron Constants (TNC) based on microscopic
 experiments results in a higher thermal fission (~587 barns) and capture (~100 barns) cross sections, and lower total nubar (2.4216 vs 2.4368)

 new resonance parameters fitted to the new TNC and resonance integrals recommended by the Neutron Standard committee, and tuned on integral data

- new evaluation of PFNS that agrees well with microscopic measurements and corresponds to a lower 235U thermal PFNS average energy (now 2.00 MeV, before - 2.03 MeV) that increases criticality





# 235U iron evaluation advances - IAEA, ORNL, IRNS, LANL & *Performance in Criticality Simulations – Low energy*

- the higher 16On, a cross sections leads to more neutron absorption and reduces criticality of epithermal assemblies

- new TSL on hydrogen slightly increases criticality of thermal solutions

- strong energy dependence of resonance nubar based on measured data from 0.3 eV up to 60 eV reduces

criticality of epithermal assemblies

- decrease of neutron capture around 1 keV and above that increases criticality for intermediate assemblies (still to be compensated)





### 235U iron evaluation advances - IAEA, ORNL, IRNS, LANL & Performance in Criticality Simulations – Fast energies

- Optical model for fission describes Neutron Standard fission cross sections on U-235 and U-238 within 3% using double and triple humped barriers with absorption. Such accuracy allows for a better prediction of elastic and inelastic neutron scattering

- Interference of direct and compound mechanisms leads to the increase of inelastic cross sections from ~50 up to 500 keV on U-238

- Neutron scattering to collective states in the continuum is important to describe observed neutron emission spectra for incident neutron energies above 1 MeV

- increased inelastic scattering on U-238 above 500 keV up to 2 MeV, slightly decreased below.





### **CIELO progress:**

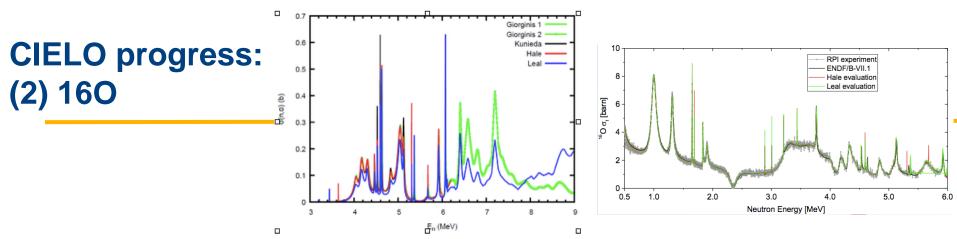
### (1) Resonance Region, extended to higher energies

Leal, Schillebeeckx, Noguerre, ... have made significant advances in representing our understanding of

- 235U resonances, notably (n,g) near 1 keV, based on recent consistent data between RPI and LANSCE/DANCE
  - updates may be needed in 10s of KeV region too
- 238U Geel measurements, which are leading to a new evaluation that is only a small perturbation compared to previous ENDF. Up to 20 keV.
  - The new (n,g) is leading to an update to the standards (similar result)
- 239Pu resonances from SG34
- 56Fe resonances up to 2 MeV with more rigorous angular distribution treatment







Various files for testing, including R-matrix analyses from Hale and from Leal

- it appears that acceptable integral performance may be maintainable, following small updates to 235U nubar, and thermal PFNS

Questions remain on the magnitude of 16O(n,alpha), with discrepancies of order 30-50%, which have ~ 100 cpm impacts on criticality.

- new measurement is planned at LANL, in Fall-2015

Low-energy thermal elastic/total - consensus reached (3.765 barns)

A new total cross section has been obtained in the few-MeV region from RPI:

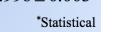
"Game changer" (Lubitz).



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3.2 MeV < E < 6	MeV C/E	C/E Statistics	Norma
ENDF/B-VII.1	0.98	8 ±0.002	$\sigma_{ ext{exp}}^{H}$
Leal 1	1.03	0 ±0.002	$\frac{v_{exp}}{u}$
Leal 2	1.00	6 ±0.002	$\sigma_{ENDF}^{H}$
Hale	1.01	2 ±0.002	2
Cierjacks 80	0.96	8 ±0.002	2

Normalization uncertainty:  $\frac{\sigma_{exp}^{H}}{T} = 0.996 \pm 0.003^{*}$ 





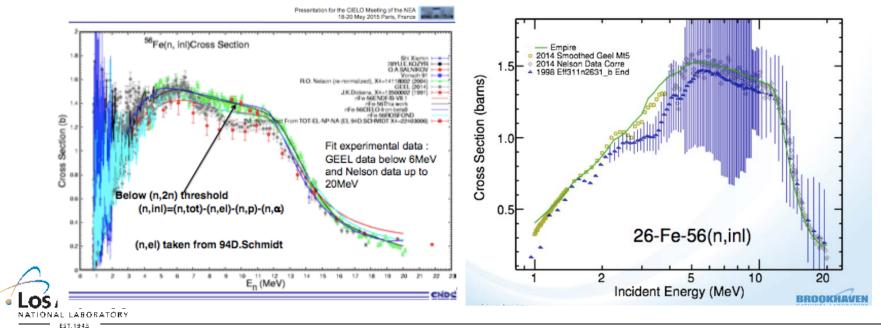
#### CIELO progress: (3) 56Fe - using advanced EMPIRE modeling, & SAMMY to 2 MeV

New 56Fe evaluation produced as a "starter file", BNL, IAEA, ... taking into account recent data from Geel & LANSCE on inelastic scattering, & RPI

- Geel data below 6 MeV accurate and consistent with new evals

Performance in iron benchmarks ~ acceptable/good ("no worse!")

Independent evaluation studies from China tend to corroborate conclusions

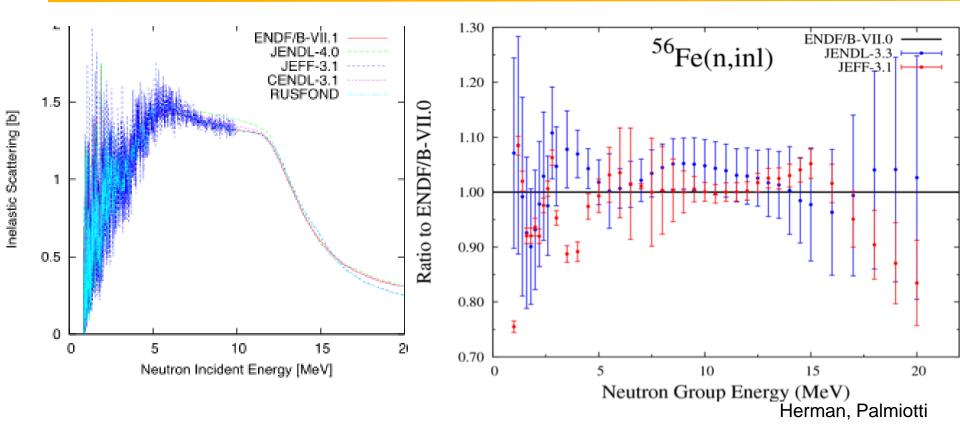




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WPEC meeting, NEA/OECD, Paris, May 9, 2016

## <sup>56</sup>Fe: Advances Needed in Inelastic Scattering



New measurements (IRMM) & SAMMY analyses in resonance region; new Hauser-Feshbach analyses at higher energies

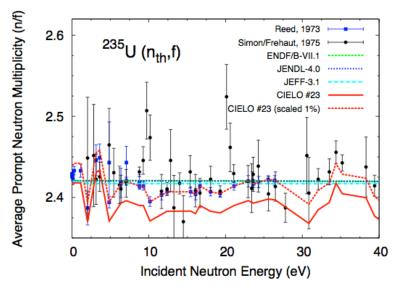
EST. 1943



#### 235 nubar? iaea

#### **Prompt Neutron Multiplicity**

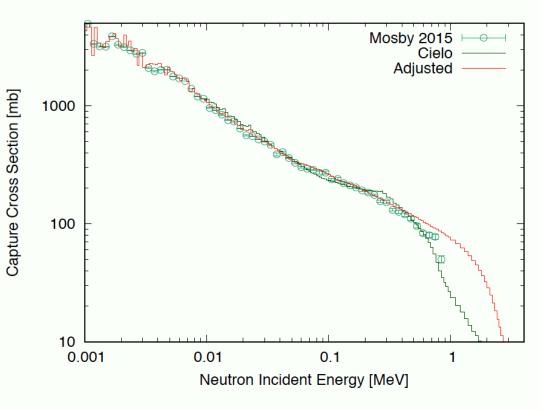
- From IAEA's latest file:
  - "Prompt nu-bar evaluated by V. Pronyaev, based on Reed data from C00-3058-39, normalized to thermal value of 2,41161.
     Simon data were measured rel. to 252Cf(sf) normalized at 10-27 eV on Reed's data (10-Jan-2016). The dip in the data around 30 eV was suppressed.
     Thermal value was increased by 0.0046 below 0.2 eV."



#### 239pu n,g

# DANCE Capture Data Test

- Mosby's data Oct. 2015
  - from URR to 1 MeV
  - URR parameters slightly adjusted
  - CoH3 calculation given to the high energy part
    - Soukhovitskii 2005 potential
    - · fission adjusted
    - M1 scissors mode included
- CIELO file issue
  - Inelastic scattering exists in URR, but total does not have it
  - Cross section fixed from 9 to 30 keV



WPEC meeting, NEA/OECD, Paris, May 9, 2016



CIELO collaboration is making progress in resolving open questions

Starter files created – being tested

Documentation to be completed in next year.

#### **Possible CIELO papers – NDS 2018?**

- Chadwick short overview of CIELO SG40 objectives and accomplishments
- Hale et al 16O evaluation ?
- Hale, Kunieda, Livermore collaborators, ...
- Broader O16 collaboration conclusions on magitude of n,a based on unitarity
- Georginis, Plompen et al conclusions on 16O (n,a) corrections on historic data
- Plompen et al, the low energy 16O total elastic cross section
- Leal? his R-matrix 16O evaluation (not in B-VIII)
- Herman et al BNL staff more details on 56Fe in fast region?
- +Trkov et al resonance region of 56Fe
- Leal et al new resonance analysis of 56Fe
- +Chinese work on 56Fe eg inelastic scattering evaluation

#### **Possible CIELO papers – NDS 2018?**

Romain, Morillon, Bauge et al - CEA fast actinide evaluations

+Chinese work on 235U

₄M.Pigni et al (Trkov) resonance analysis of 235U

- R. Capote et al, fast 235U and 238U analysis
- ↓L. Leal resonance analysis of 235U

+Schillebeeckx et al., new Geel 238U resonance analysis - MBC notes not yet in B-VIII-beta1

Kawano - 238U and LANSCE capture theory and moel calculations

Neudecker - PFNS evaluations of actinides - incl those that did not make it into B-VIII, eg ther mal 239Pu

+Talou - multiplicity dependent fission neutrons and gamma-rays

Kawano, Capote, Romain, et al, latest conclusions on actinide inelastic cross sections

#### **CIELO NDS 2018 Template**

#### Neutron reaction on XXX under the CIELO Collaboration

author<sup>1,\*</sup>

<sup>1</sup>Lab address (Dated: April 29, 2016)

Abstract.

Note we do not yet know page restrictions for these individual CIELO papers, but plan for somewhere between 4-10 pages each. (The same ND2018 issue will have a large 100+ paper on ENDF/B-VIII, and a large 30+ page paper on new standards). Thus we expect perhaps 150 pages reserved for separate CIELO papers.

#### I. INTRODUCTION

The paper should be journal quality; it will be peer reviewed.

**Give background, e.g.** - The CIELO pilot project was commissioned by the OECD's Nuclear Energy Agency WPEC (Working Party on International Nuclear Data Evaluation Co-operation) during a meeting held in May 2012. The goal has been to identify deficiencies and discrepancies in our current understanding of neutron reactions on high priority nuclides <sup>1</sup>H, <sup>16</sup>O, <sup>56</sup>Fe, <sup>235</sup>U, <sup>238</sup>U and <sup>239</sup>Pu, and to develop proposed solutions and improvements in our understanding. The goals of CIELO are documented in Ref. [?]. This reference, together with other papers such as Refs. [???] document some of the questions being addressed.

Explain clearly which isotope and cross sections or energy/angle distributions will be addressed.

Explain the background as to why our previous understanding was inadequate. Why - a lack of accurate experimental data? Insufficiently reliable model predictions? What were the assessed uncertainties prior to the present work?

Possibly summarize the applications that have motivated the work.

#### II. RESULTS

Explain the basis for your new conclusions - new experiments made, theory, new analyses?

What are your new proposed uncertainties?

Proposed or recommended evalations should (a) provide a ENDF-formatted file that can be archived on the NEA CIELO web site; and (b) be documented in this paper, with illustrative figures that compare the recommendation against the main evaluated data libraries, such a ENDF/B-VII.1 [? ? ], JEFF-3.1 [? ? ], JENDL-4.0 [? ], BROND/ROSFOND [? ], and CENDL-3.1 [? ].

What are remaining open questions that remain unsolved?.

#### **BACKUP SLIDES**

WPEC meeting, NEA/OECD, Paris, May 9, 2016

# Cecil Lubitz:

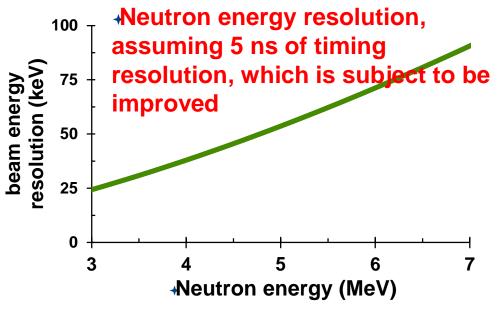
"After several "preliminary" months on CIELO it's clear that we have bitten off a big chunk. Get ready to chew."





# LANL plans to measure new cross sections on <sup>16</sup>O(n,( )

- OUse a newly developed instrument LENZ with a large solid angle and low alpha detection threshold
- OUse a white neutron source at WNR/LANSCE in Fall 2015
- **ORelative measurement to <sup>6</sup>Li(n,a) reaction to reduce systematic uncertainty**
- **OFirst goal is to measure cross sections at the energies between 3 and 5 MeV**

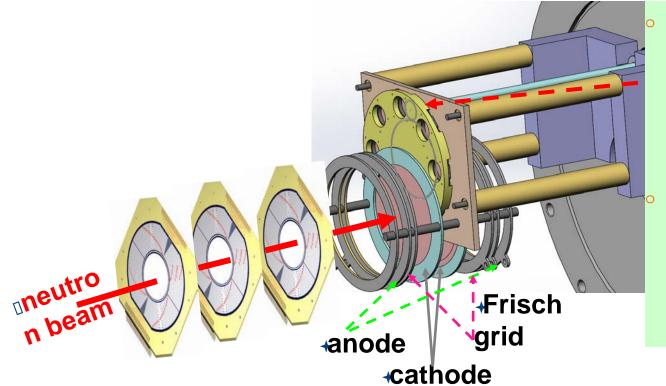




Picture of LENZ chamber

WPEC meeting, NEA/OECD, Paris, May 9, 2016

# LENZ : Twin Frisch-grid Ionization Chamber + Silicon Strip Detectors



The multiple target system allows to have a oxygen and a Li reference target at the same time Solid oxygen target is made by anodizing highly-

enriched water on

tantalum backing

Cathode OAt forward angles, the silicon strip detector measures angles and charged particles energy deposit

ODigitizers provide wavelet information for post processing of improvement of signal-to-noise ratio with no dead time

### 16**O**

# Hale comments on new (n,tot) RPI data:

Comments about LANL n+16O Cielo evaluation:

The evaluation we submitted in June of last year is similar in many ways to ENDF/B VI.8. For that reason, since the total cross section was preserved in the evaluation that finally became ENDF/B VII.1, it is not surprising that the agreement with the new RPI measurement of the total cross section looks similar for VII.1 and the LANL Cielo file. Our latest evaluation is somewhat better in the "window", and somewhat worse at energies above about 4.5 MeV. Adding these total cross section data to the analysis would likely decrease the total cross section somewhat in the 2.5-3.5 MeV region, which because of the often-noted anti-correlation effect of unitarity in this region, would tend to raise the fitted (n,alpha) cross section at these energies. This would make the disagreement even worse with experiments that favor the lower normalization scale for the reaction cross section.

We are anxious to add these measurements (not the binned data) to our analysis to see what their effect might be, but we are gratified that the initial comparison does not seem to indicate any major problems with the evaluation. Hopefully, we will have additional (n,alpha) data coming from Los Alamos in the next year or so. In the meantime, we are working on extending the existing LANL file above 6 MeV, and including the Geel (n,alpha) data in our analysis, following Giorginis' recommendations about normalizations, etc.





2			
3	160		
4			
5			
6			
	General		
8		Intercompare evaluations, and identify goals for a new evaluation	
9			
10		JENDL is a new work (though adopts ENDF n,a); ENDF (JEFF uses ENDF) is a hydrid of KAPL work < 3.2 MeV, LANL (Hale et al) > 3.2 MeV - assess value	
11		The 2005 ORNL work generated a resonance analysis for 160, full R-matrix. Included angular distributions, n,alpha, and it has never been tested. Neede	ed L
12			
13			
4	Tabal Theatle		
	lotal, Elastic	A forward and the scale of the	
16 17		Compare existing evaluations and R-matrix analysis, and defie path forward	
18		At low energies, assess whether evaluations of elastic scattering indeed need to be lowered by $\sim$ 3%, as proposed by Plompen, Lubitz, Roubtsov etc	
19		At low energies, assess whether evaluations of elastic scattering indeed need to be lowered by ~5%, as proposed by Plompen, Edbizz, Roubisov etc	
20		covariances for mubar: Need reliable anisotopic 160 scattering uncertainties. Palmiotti thinks Gerry's present uncertainties are too small on mubar.	
21			
-	Capture	ENDF adopted JENDL's capture cross section to include resonance contribution - establish consensus to use this	
23			
24			
	(n,a)		
26		Review different evaluations (all largely same as ENDF)	
27		Review previous data, and agee on scales - eg Bair & Haas had renorm their original data down by ~20%; Are Johnson data the same as these?	
28			
29 30		Review new data - Georginis (Geel), Khryachkov (IPPE) - contact physicists working on 13C(a,n) for astrophysics	
30 31		The above new data approx confirm ENDF below 6 MeV but point to changes above Intercompare R-matrix calcs (Hale, Kunieda, Leal)	
32		Seek to understand why the above R-matrix evaluations, influenced by total cross sec data, suggest ~30% higher (n,a) than most measurements	
33		Define an evaluation strategy If theory contradicts these data, do we use data instead? Or do we conclude theory is right and measurements had a scle	e er
34		Assess whether evaluations (all now based n ENDF) above ~ 6 MeV need changing, if it is concluded new Geel data are more accurate than old Davis da	
35			
	Integral	Establish suite of integral valdation tests, including k-eff, transmission, etc	
37		2 benchmarks sensitive to oxygen data (+11 more benchmarks with water) are available in the SINBAD database	
88		Broomstick experiment	
39		Following WPEC SG?, With the existing (n,a) evaluations perform well, for the most part, on LEU solutions, Can the new eval perform well too	
10		(n,a) impact at higher energies: Does this higher energy >6 MeV region impact any applications significantly (maybe medical applications)? Carlson note	es M
11		check astrophysics constraints on 13C(a,n) reaction rate	
12			
13 14			
14 15			
+5 46			
17	E		



### 1H & Other Standards: Hale's Summary and Outlook

- NN analysis progressing; more p-p elastic scattering data needed in the 30-50 MeV range. Low-energy parameters retain their earlier (correct) values. Need to extend analysis above 200 MeV.
- New data for n+<sup>6</sup>Li fit in well with the existing data set, and cause no problems with the R-matrix fitting.
- n+<sup>12,13</sup>C analyses in good shape below 2 MeV. Could produce a natural C standards file in this energy region now. More work is needed on both evaluations at higher energies, however.
- Problem with unrealistically small uncertainties on standards cross sections may be solved by using parameter confidence intervals.

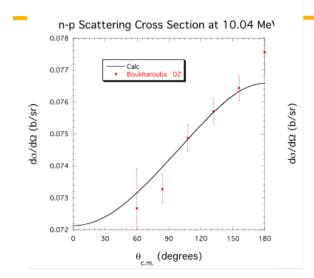
### <sup>1</sup>H – Hale comments on the covariance data

The uncertainties in the n-p scattering cross sections that were put into VII.1 (as described in my CW 2008 paper) are fairly realistic (maximum of 1% at around 10 MeV). The uncertainties on the capture cross section are probably too large, due to the kludge I had to make to compensate for Lubitz's insistence that the thermal value be a certain number. All of this should be better in the next release, since we will use confidence intervals in place of standard deviations (which has the effect of scaling up the standard deviation by a known amount)



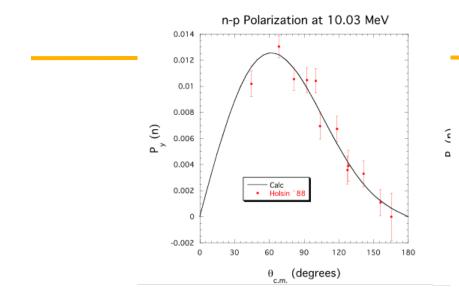


### n-p Differential Cross Sections



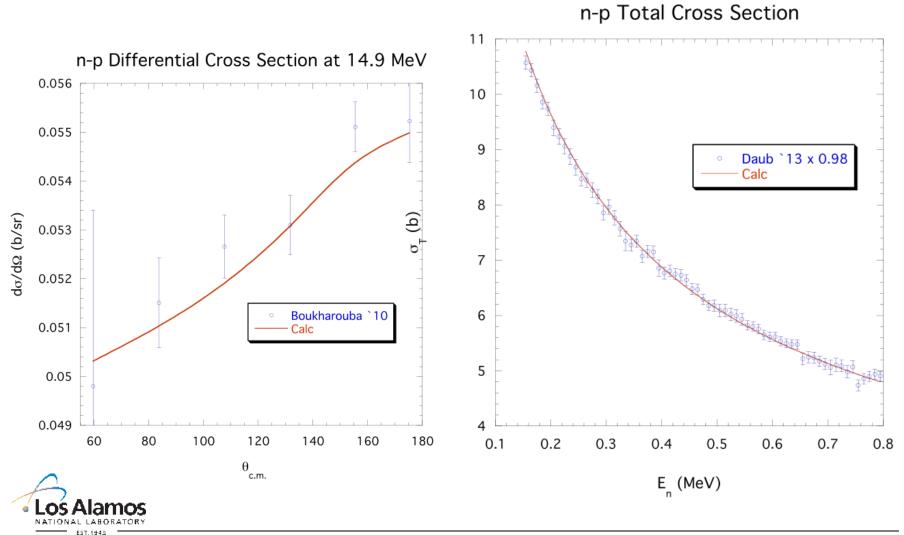


## **n-p Polarizations**





### <sup>11</sup>H recent data added to analysis



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WPEC meeting, NEA/OECD, Paris, May 9, 2016



4	A B	C	D
CIELO	: Summary of tasks to address:		
2			
Act	inides: 239Pu, 235U, and 238U - specific issues for each nuclide a	re noted	
Fact	Region (keVs and above to 20 MeV) - fission listed separately		
Tast	Region (Revs and above to 20 MeV) - insitin insted separately		
Boulou	v Overall Goals, as embodied in this document and in LAUR CIELO document		
Review	Voverali Goals, as embodied in this document and in EAOK CIELO document		
Inelast	tic and elastic scattering - below a few MeV (eg 7)		
)	Review existing discrepancies between evaluations		
	Collect all available experimental data		
2	Review various theoretical approaches, as embodied in codes (including HF, Coupled Channels, KKM,)		
- 3	Discuss and review optical model options		
4	238U: dispersive coupled-channels OM developed at IAEA		
5	Seek consensus on best evaluated reprentation of data		
5	238U: 238U Elastic and inelastic scattering data from RPI. Quasi differential available (mainly inelastic) from	RPI from from 0.5 MeV u	n to 20 MeV - F
7	235U: New (n,xng) data to be published in PRC by Kerveno et al. (IPHC, Strasbourg (F)) could be useful to r		
8	Understand implications from integral data testing on changes in inelastic scattering - especially k-eff and re		
9	Assess covariances and implement in ENDF format	decion nates (speed an indi	
0	Create ENDF formatted files		
	tic and elastic scattering - 7-20 MeV		
2	Review existing discrepancies between evaluations, data, and models (including preequilibrium)		
3	Collect all available experimental data - including Kammerdiener's data and Baba's (U8) data		
4	Review various theoretical approaches, as embodied in codes (including preeq, HF, Coupled Channels, KKM,	PFNS background,)	
5	Discuss and review optical model options	;,,,	
6	Seek consensus on best evaluated reprentation of data - including possible continued use of pseudostates		
7	Understand implications from integral data testing on changes in inelastic scattering -especially 14 MeV puls	ed spheres/transmission of	iata
8	Assess covariances and implement in ENDF format		
9	Create ENDF formatted files		
	on Capture		
1	239Pu: Review discrepancies between evaluations, which exceed 10% at the higher energies		
2	235U: Review discrepancies between evaluations, which exceed 25% near 1 KeV (Japan'shigher result) and	10% at the higher energie	es
3	238U: Consider adopting 238U capture from standards - ENDF/B-VII used this, but with some small differen		
4	238U: Monitor Standards results for any changes, based on new measurements from DANCE, nTOF, Geel		
5	239Pu: Review data (very few measurements, especially above 100 keV there is just the LANL Hopkins data	); See if DANCE data is av	ailable in time
6	235U: Review new DANCE data and RPI data, that appear to corroborate JENDL changes near 1 keV, but pol		
7	Review guidance from integral PROFIL data (suggests PU9 and (maybe) U5 from ENDF should be higher), and	nd Wallner AMS data at 25	keV and 420 ke
8	Assess model calculations predictions (consisent with above inelastic scattering HF/CC/OM calculations)		
9	Seek consensus on best evaluated reprentation of data		
0	Understand implications from integral data testing on changes in capture - especially k-eff and reaction rates	s (spectral indices for 85/5	of etc)
1	Assess covariances and implement in ENDF format		
2	Create ENDF formatted files		
3			
4 n2n			
5	Discuss data, including discrepancies in rise from threshold, and differences near 14 MeV		
6	Review existing evaluations (including "GEANIE evaluation" for 239Pu), data , and calculation predictions		
7	235U: New (n,xng) data to be published in PRC by Kerveno et al. (IPHC, Strasbourg (F)) could be useful to r	model n2n scattering, see	prelim results in
8	239Pu: Carefully note insights on n2n making 238Pu from LANL, and discuss contradictory feedback from PF	ROFIL	
9	Validate any changes against n,2n reaction rates in critical assemblies, eg Fig 57 in NDS112,(2012) ENDF		
0	Create ENDF file and covariances		

10				
71	Fission (all	energies), cross sections, nubar and spectra for n,g		
72				
	Review Overall (	Goals, as embodied in this document and in LAUR CIELO document		
	Fission Cross Se			
75	1			
76	,	Seek consensus that we adopt the fission cross section standard from the IAEA group		
77		Assess implications of adopting standard fission cross section on integral testing		
78		If IAEA standards team updates their value, use it; this would include any recent/forthcoming fission measurements, eg	nTOF, RPI, TPO	С
79		Modeling of fission would occur as part of the above inelastic/capture/n2n activities, but seek consensus that we do not		
80	1			
81	1	238U:Subthreshold fission for 238U - discrepacies between different evaluations. Lead spectrometer measurements nea	ar 70 keV sugg	est a p
82	prompt nubar			
83		Review existing evaluations and experimental data, & review various theoretical approaches; 238U low energy interp fix	needed in END	<b>DF</b>
84		Seek to use an "unadjusted" nbar in a final evaluation, avoiding the ENDF "tweal" near an MeV that was adopted to bett		
85		Study Koning-Rochman nubar near thermal, from their optimization search (but it's 3 SD below the standards constants		
86		Develop a new evaluation based on a covariance analysis of the data		
87		Understand implications from integral data testing on changes in nubar - especially k-eff		
88		Create ENDF formatted files, including covariances		
89	,			
90	PFNS	Review work of IAEA CRP on PFNS		
91		Aim to adopt the CRP's recommendation		
92		Seek consensus on using LANL high-accuracy NUEX Pu9 and U5 data, as published in Dec NDS2011 to help define high-	energy spectru	m
93		Use new PFNS measurements, especially below MeV, coming from LANSCE/Chi-nu in the coming years		
94		Use guidance on high energy tail of spectrum from dosimetry reactions (new IAEA IRDFF CRP), eg from LANL crits, Russ	sian fast reacto	r, & CE
95		As part of IAEA CRP, advance our theoretical models, and use incorporate other data (new and existing)		
96		Understand implications from integral data testing on changes in inelastic scattering - especially k-eff and reaction rates	in assemblies	
97		Create ENDF formatted files, including covariances		
98	,			
99	,			
	PFGS			
101		Review existing evaluations and experimental data, and various theoretical approaches		
102		Represent fission gammas separately at all energies, including above 1.09 MeV for U5 and Pu9 (an ENDF drawback), & u	use new data a	vailabl
103		Update PFGS spectra to use modern measurements from DANCE, as well as multiplicity distribution if possible		
104		Create ENDF formatted files, including covariances		
105	,			
106	Delayed data			
107		Review differences in present evaluatiosn		
108		Develop plan for work needed		
109	,			
110	Energy Release			
111		Compare energy release data in evaluations, for prompt n, g, fission fragments; and delayed energy release		
112		Update as necessary - eg ~ MeV level changes are impled for 239Pu from Jandel's DANCE data for 239Pu (but 235U lool	ks good)	
113		Consider updating energy release incident-energy-depenence based on Lestone's work		
114				
		the concerning here ceep, rand, may be re-		·

110					_
116	Integral Da	ta Testing and Validation			
117					
118	Review Overall	Goals, as embodied in this document and in LAUR CIELO document			
119		Define suite of critical assembly, reactor, transmission, etc experiments to use in validation assessments, and observable	es (k-eff, rates,	spectral indic	ces)
120		238U: selection of 12 ICSBEP criticality benchmarks sensitive to elastic scattering is available from JSI/IAEA (Trkov, Cap	ote)		
121		Seek to ensure good performance in data testing, which includes:			
122		Fast, intermediate, and thermal assemblies, k-eff			
123		239Pu: Aim for (Partial?) improvement of longstanding overprediction of thermal Pu solutions			
124		Modeling spectral indices well in various systems (incl fast), 8f/5f, 9f,5f, 237np-f/5f, 233u-f/5f etc, see Table XXXVIII in	VII.1 NDS 2011	paper	
125		Modeling of post irradiation experiments (PIE) such as PROFIL (CEA) and MANTRA (INL)			
126		Modeling MOX experiments for mock up of LWR, eg in EOLE, Cadarache			
127		See if PFNS improvements give improved n2n detector responses in fast crits, eg through a softer PFNS spec aove 10 M	eV		
128		nubar validation using multiplication subcritical measurements			
129		LLNL pulsed spheres			
130		Can we obtain improved preductions of intermediate assemblies, eg ZPR at Argonne			
131		Aim to maintain good prediction of crits, including new as-built high-resolution 3D MCNP Jezebel model?			
132		Use sensitivity metodologies for assessing changes/improvements by reaction and energy range			
133					
134					
135					
136					
127					

2						1	1				
3	56Fe										
4											
5											
6											
7	General										
8		Review differer	nces in evaluation	ons. In ENDF/B	-VII.1 RR exter	nd up to 850 k	eV, but pointwis	e fluctuations e	xtend up to alm	nost 10 MeV.	
9							rkov, Koning, Vo				bean Jeff ev
10			and other key n								
11											
12	Fast Region										
13	. act ricg.co.										
14	Inelastic and ela	stic									
15	inclustic and cit		ta.: RPI has hid	h-res transmis	sion up to 2 M	eV. and scatter	ing data ("quasi	differential dat	a"), that needs	an MCNP calc t	to compare
16									,,		
17		Review new da	ta:Arian Plomp	en (Geel) has i	nelastic data (a	actually, gamm	a-production) to	o measured th	s vear, from 80	0 keV to 5 MeV	i.
18							measurements				
19			ta: Ron Nelson								
20							is suggesting a	big change for	nonelastic, but	that our total	cross sectio
21			channel OM wor								
22		Pronyaev - als	o doing work or	n inelastic gami	ma production.	At one point t	his was being co	onsidered as a s	tandard (now r	nore likely to u	se Ti).
23											
24	Charged-particle	e production									
25			valuations, and								
26		Data above 20	MeV may be no	eeded too, eg fo	or fusion applic	ations, using n	ew gas-product	ion data from H	laight.		
27											
28	Activation xs										
29		Review/Include	e activation dat	a needed for fis	sion/fusion						
30											
31	DPA	Take advantag	e of insights fro	m new IAEA CF	RP on damage	and DPA					
32											
33	Resonance	Region, Re	solved and	UnResolved	d Paramete	ers (hundre	d of keVs a	nd below)			
34											
35	RRR & UR	Review latest e	evaluation from	Luiz Leal							
36											
37	Integral validati										
38							eactor experime				
39							el) are available				
40							perience, etc, St		S 2012 benchm	arking paper (	which note:
41							benchmark for				
42							m CEA, e.g. PE	RLE experiment	s in EOLE)		
43			ZPR9-34, ZPR6								
44		Use sensitivity	metodologies f	or assessing ch	anges/improve	ements by read	tion and energy	range			
45											
46			,,	l	n Lo meeting,	NEALOLOD, 10	ans, way 7, 201	10	1		



# Pu-SOL-THERM Benchmarks – II. Prelim LANL testing of new Subgroup 34 resonance results

- A set of seven Pu-SOL-THERM benchmarks have been extracted from the larger set.
  - PST1.4 & PST12.13 span the ATLF space;

•

- PST12.10 & PST34.15 span the ATFF space;
- PST4.1 & PST18.6 span the <sup>239</sup>Pu atom percent space;
- PST12.10 & PST34.4 span the g Pu per liter space.
- All benchmark experiments are performed in simple geometry
- PST1.4 & PST4.1 are a water-reflected spheres;
- PST18.6, PST34.4 & PST34.15 are water-reflected cylinders;
- PST12.10 & PST12.13 are a water-reflected slabs;

# Pu-SOL-THERM Benchmarks – III. Prelim LANL testing of new Subgroup 34 resonance results

### +The E71 1.00576

k<sub>calc</sub> average demonstrates that the 7 benchmark subset reflects the larger population.

 Data revisions in the "Leal7a" <sup>239</sup>Pu evaluated file have eliminated ~50% of the longstanding k<sub>calc</sub> bias.

Assembly	ENDF/B-VII.1	JEFF-3.1.2 <sup>(b)</sup>	JENDL-4.0 <sup>(b)</sup>	Leal7a <sup>(c)</sup> + e71	Leal7a (RR, nu, pfns only) + e71
PST1.4	1.00448	1.00127	1.00588	1.00199	1.00202
PST4.1	1.00383	0.99907	1.00482	1.00044	1.00044
PST9	1.01939	1.01367	1.02510	1.01543	1.01546
PST12.10	1.00412	0.99973	1.00498	1.00083	1.00080
PST12.13	1.00955	1.00468	1.01069	1.00611	1.00620
PST18.6	1.00472	1.00153	1.00557	1.00202	1.00208
PST34.4	1.00258	0.99999	1.00417	0.99922	0.99937
PST34.15	0.99742	0.99563	0.99844	0.99679	0.99707
Average	1.00576	1.00195	1.00746	1.00285	1.00293

Using Various <sup>239</sup>Pu Cross Sections

Calculated Eigenvalues<sup>(a)</sup> for a Selection of PST Assemblies

a) MCNP calculations are for 250M histories; stochastic uncertainty is ~5 pcm.

b) JEFF-3.1.2 and JENDL-4.0<sup>239</sup>Pu only; remaining nuclides are ENDF/B-VII.1

c) "LEAL7a" evaluation provides revised resolved resonance parameters coupled to a joint ORNL/CEA evaluated <sup>239</sup>Pu file; the "LEAL7a (RR,nu,pfns)" file couples just these data to the existing ENDF/B-VII.1 <sup>239</sup>Pu file.

# **Time-line**

### May 2013: CIELO WPEC Subgroup initiated

• Teams identified

### Nov 2013: NEMEA7-CIELO: Main collaboration kick-off

- Refine scope of work, collaborators who will work on tasks
- Will result in detailed work plans, time line goals, for each nucleus

#### Next 2014-2016 Years:

- Various collaboration meetings, continual email collaborative exchanges
- engagement with validation data testers continually
- Start incorporating new IAEA standards results (fission, capture, scattering, ...)
- Explore interdependencies on criticality from the 6 CIELO nuclides

### May 2016:

- Document conclusions from CIELO collaborations in WPEC report (& NDS paper, 2018)
- Create formatted files that embody CIELO's initial conclusions



